

# MEGHA - TROPIQUES

## Scarab, Calibration Results

14/02/2014-Paris

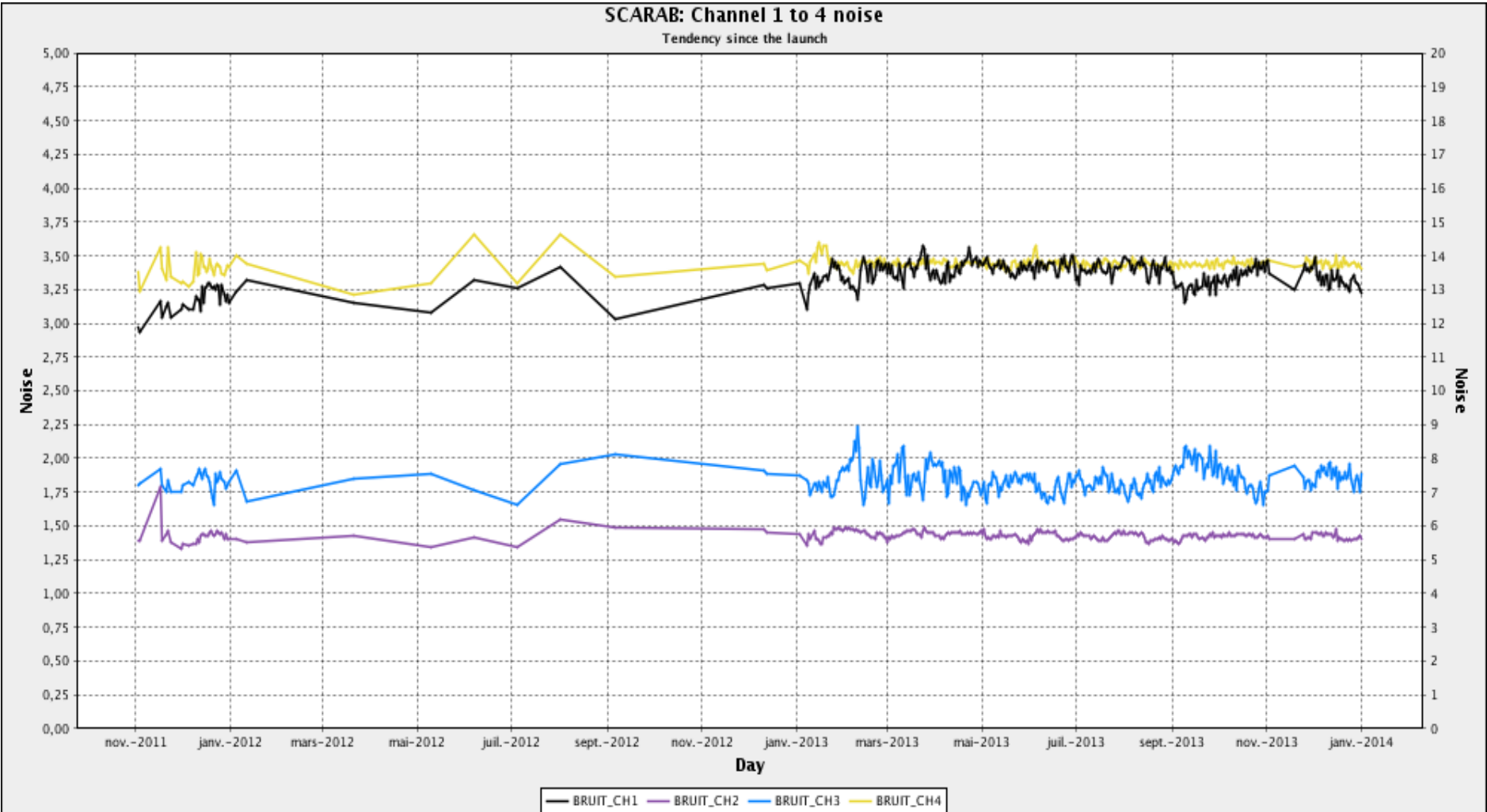
- ◆ Radiometric noise
- ◆ Gain
- ◆ A' factor

## Radiometric noise

### Requirements :

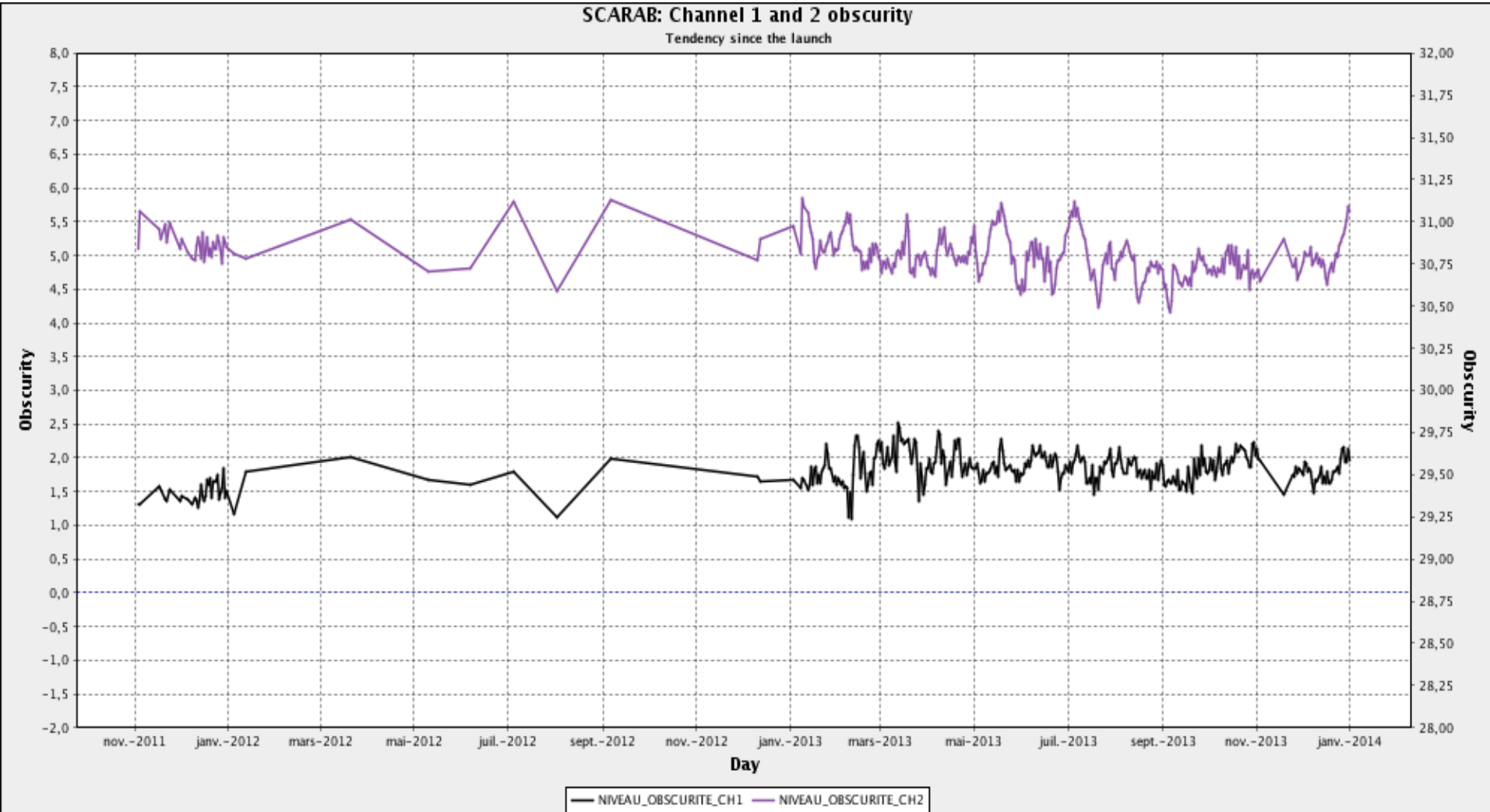
<i>Channels</i>	<i>Noise</i>	<i>Noise (LSB)</i>
<i>Channel 1 – Visible channel</i>	$< 1 \text{ W.m}^{-2}.\text{sr}^{-1}$	<i>30 LSB</i>
<i>Channel 2 – Solar channel</i>	$< 0.5 \text{ W.m}^{-2}.\text{sr}^{-1}$	<i>15 LSB</i>
<i>Channel 3 – Total channel</i>	$< 0.5 \text{ W.m}^{-2}.\text{sr}^{-1}$	<i>15 LSB</i>
<i>Channel 4 – IR Window</i>	$< 0.5 \text{ W.m}^{-2}.\text{sr}^{-1}$	<i>100 LSB</i>

## Mean Noise in LSB per orbit



=> Noise stability is very good

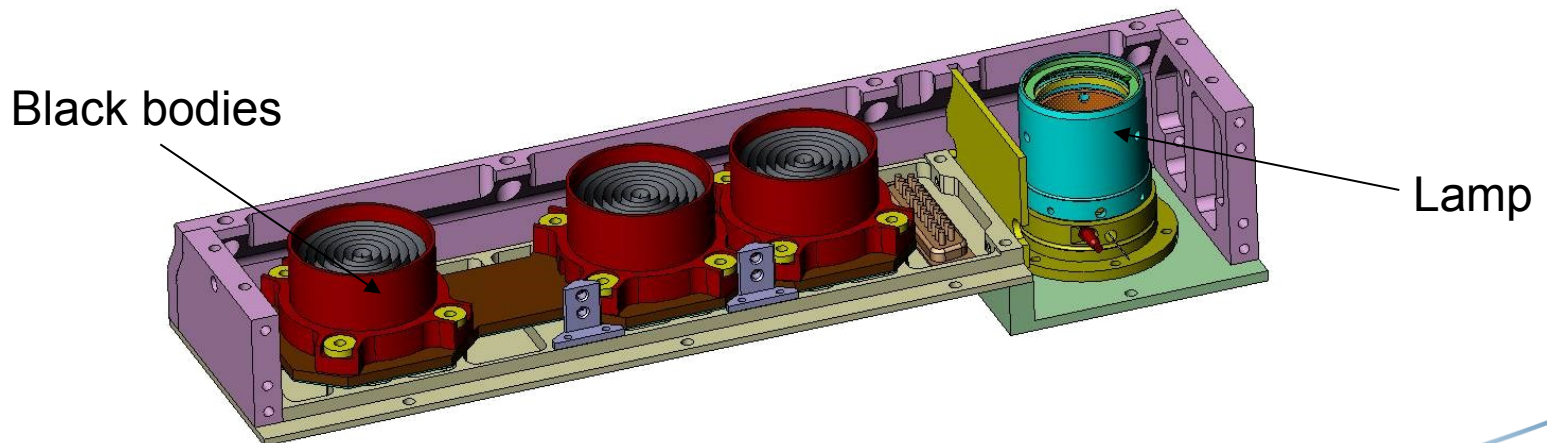
## Obscurity level for channel 1&2



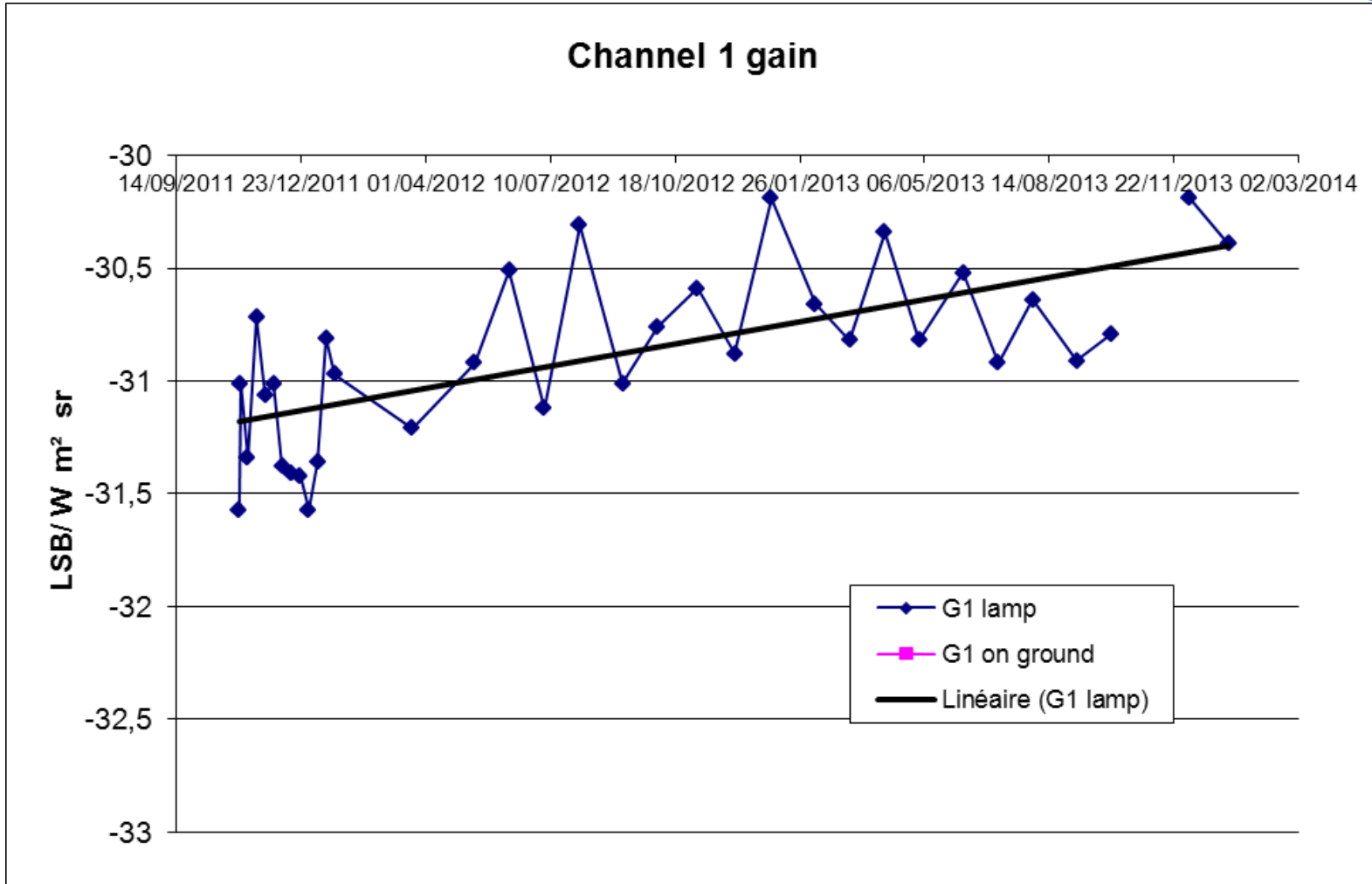
## Gain

### Measured with the CALibration Module (CALM)

- ◆ 1 lamp for channel 1
- ◆ 3 black bodies for channel 2-3-4
- ◆ No solar filter on channel 2 (filter wheel)
- ◆ CALM mode once per month



# SCARAB gain



**1% per year of drift ?**

## Gain canal 1, autre méthode

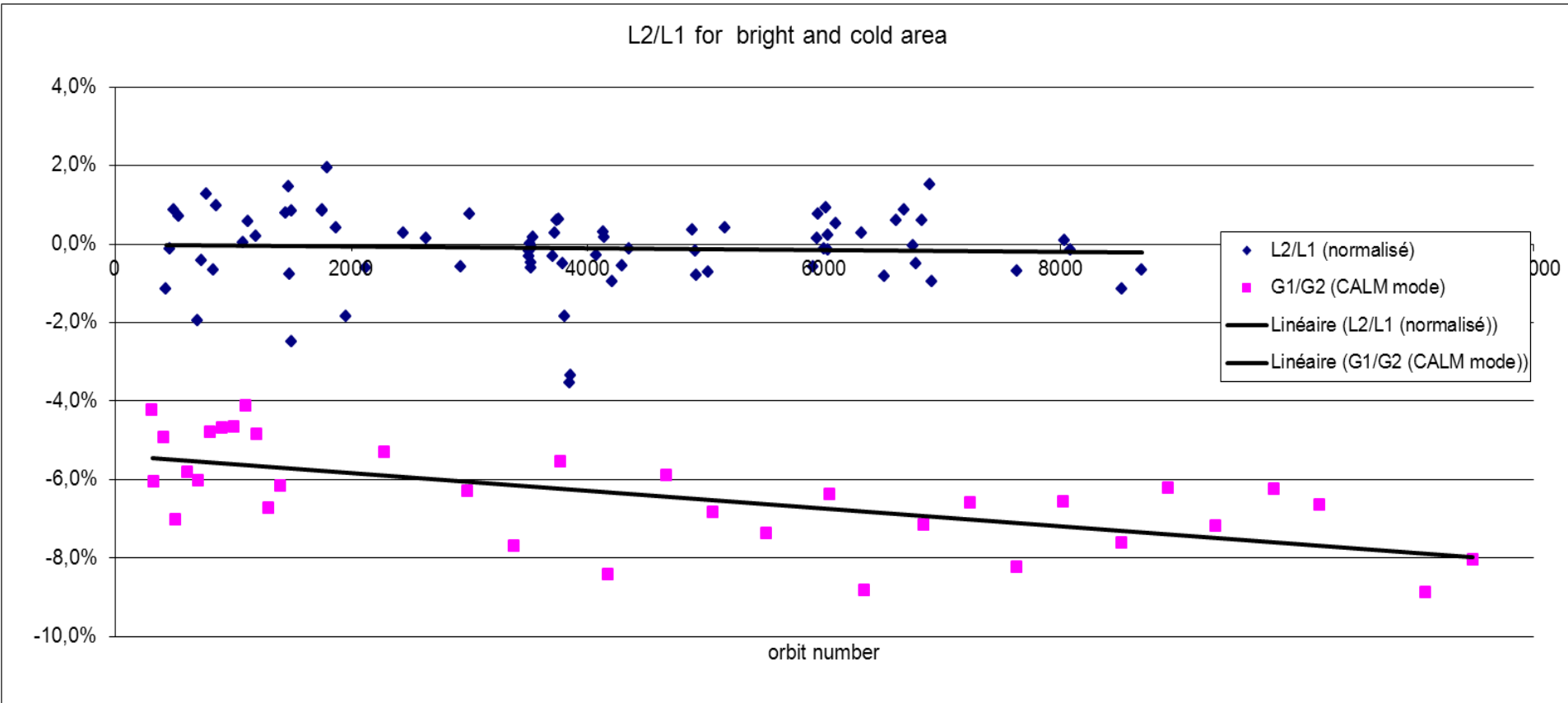
### Comparaison des comptes numériques canal1 / canal 2

Scène homogène, lumineuse et froide (nuage)

- ◆ Canal 4  $< 5\text{W/m}^2/\text{sr}$
- ◆ Canal 2  $> 250\text{W/m}^2/\text{sr}$
- ◆ Zone homogène à 10% près sur la radiométrie du canal 2.

Calcul du rapport C1/C2, normalisé

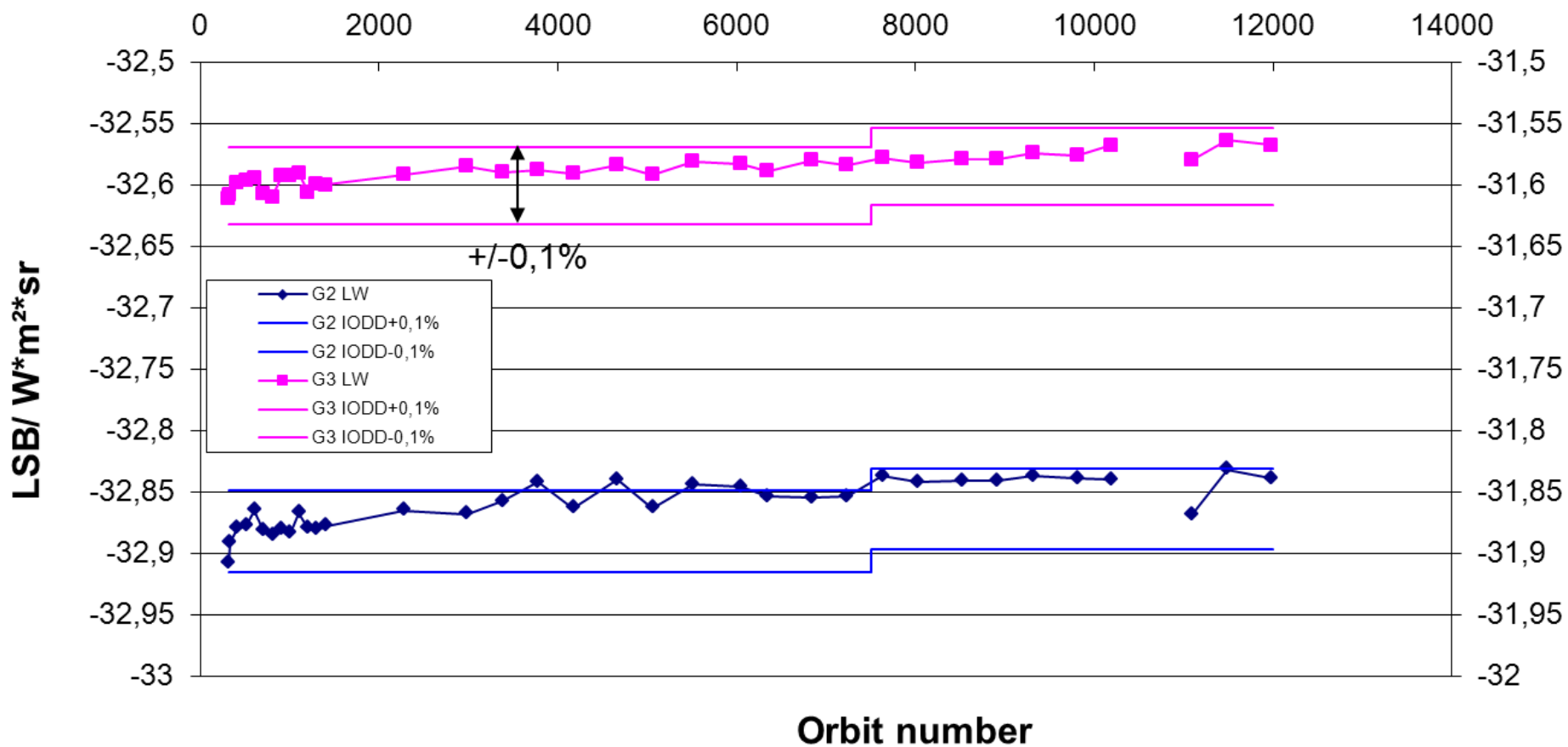




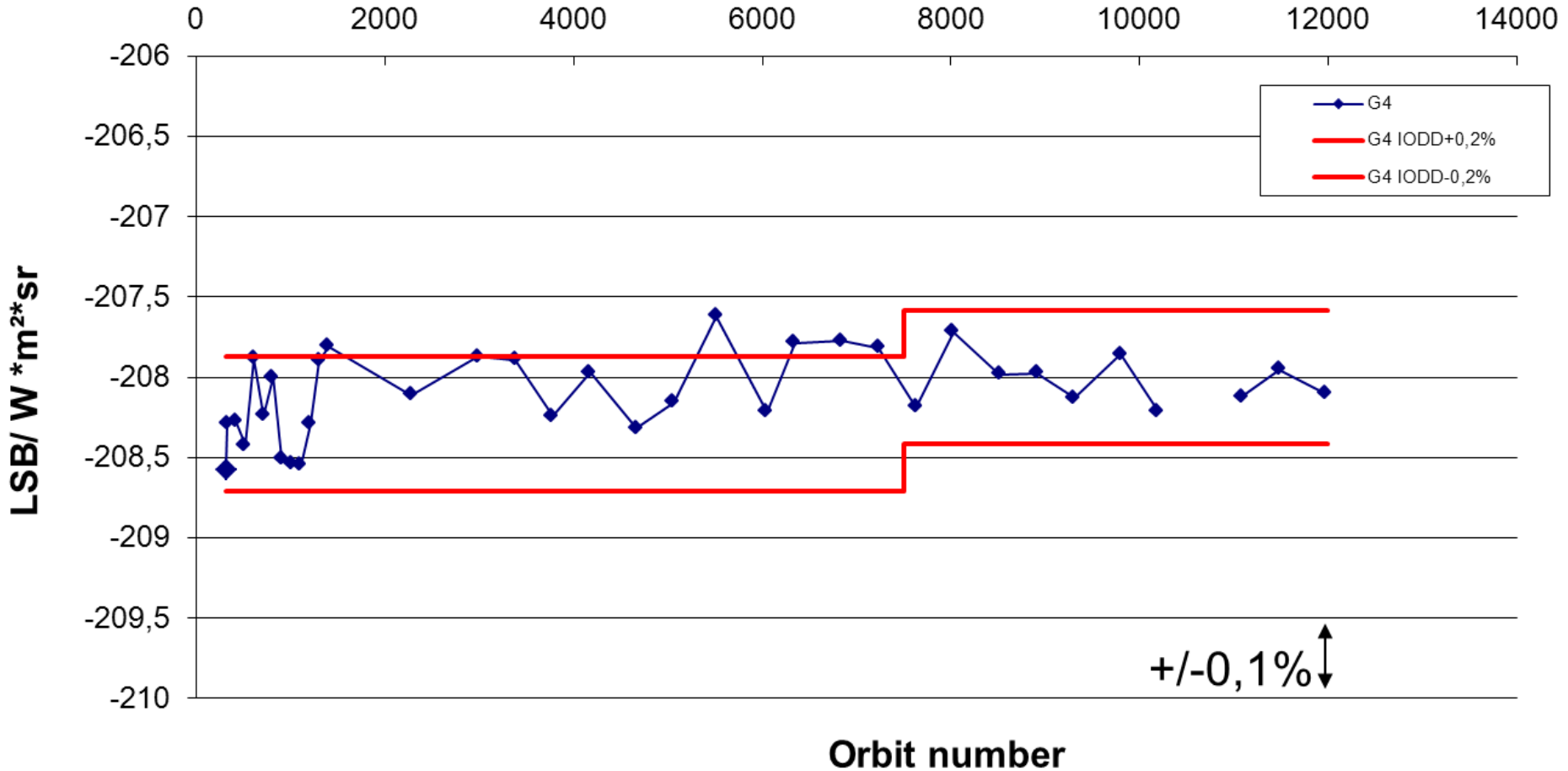
**La stabilité du canal 1 est identique à celle du canal 2**  
**La lampe est instable (1% par an environ)**

**=> Nouvelle méthode de suivi du canal 1 à implémenter**

## LW Gain for Scarab channels 2&3



## LW Gain for Scarab channel 4



## Gain

**Gain stability is for channel 1 is similar to channel 2**

**Gain stability is around -0,1% for 2 years for channels 2&3.  
=> Slight update to be done soon (+0,05%)**

**Gain stability is better than +/-0.2% for channel 4.**

**Gain will be probably slightly updated soon**

<b><u>Channel 2</u> :</b>	<b>Solar channel</b>	<b>0.2-4<math>\mu</math>m</b>	<b>Short wave radiance</b>	<b>L<sub>sw</sub></b>
<b><u>Channel 3</u> :</b>	<b>Total channel</b>	<b>0.2-200<math>\mu</math>m</b>	<b>Total radiance</b>	<b>L<sub>total</sub></b>
<b><u>Channel 5</u> :</b>	<b>Infrared channel</b>	<b>4-200<math>\mu</math>m</b>	<b>long wave radiance</b>	<b>L<sub>lw</sub></b>

**Channel 5 is computed with :**

$$L_{lw} = L_{total} - A' \times L_{sw}$$

**When channels 2&3 observe a same pure SW source, A' can be evaluated by :**

$$A' = L_{total} / L_{sw} = L_3 / L_2$$

**A' represents the difference of sensibility in the SW domain, between channel 2 and channel 3.**

## MS Mode :

Channel 2 AND Channel 3 have an identical silica filter

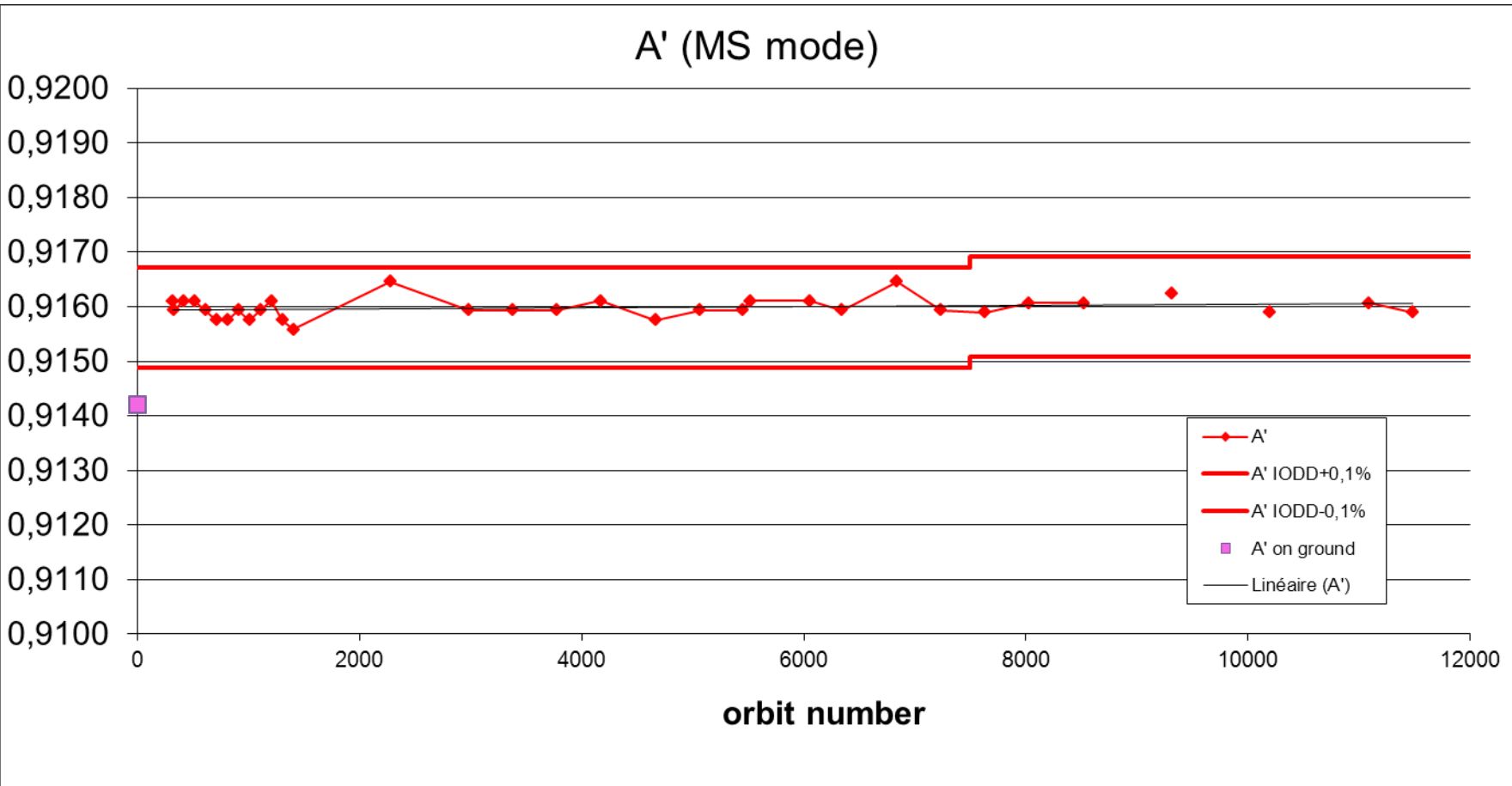
=> Both observe the same pure SW source

$$\Rightarrow A_{ms} = N_3 / N_2$$

$$\Rightarrow A' = A_{ms} * G_{2\_sw} / G_3 / T_{filter}$$

$T_{filter}$  must be known accurately with on-ground measurements !

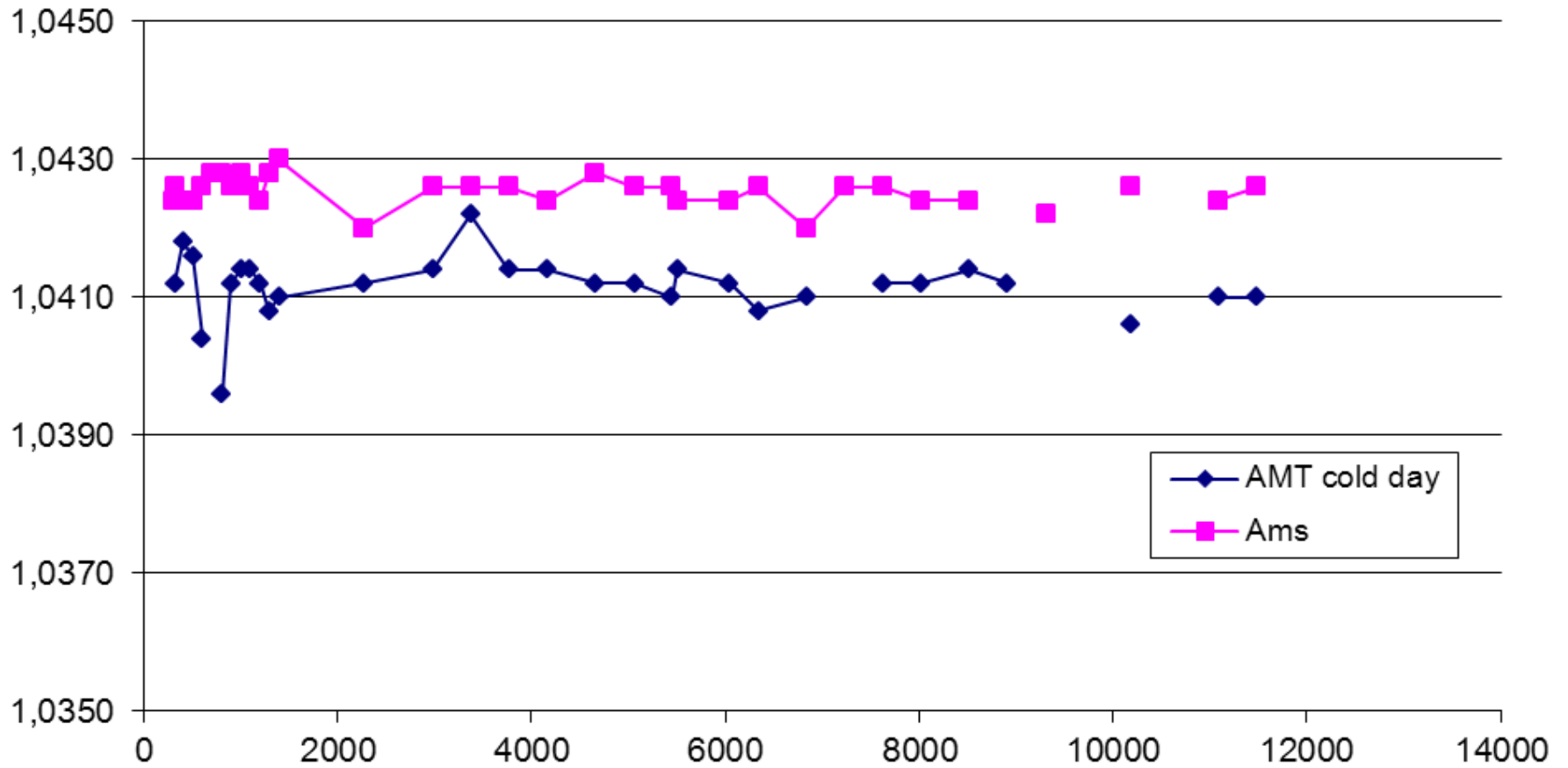
## MS Mode :



The stability of the A' factor is about +/-0.05% for these two years !

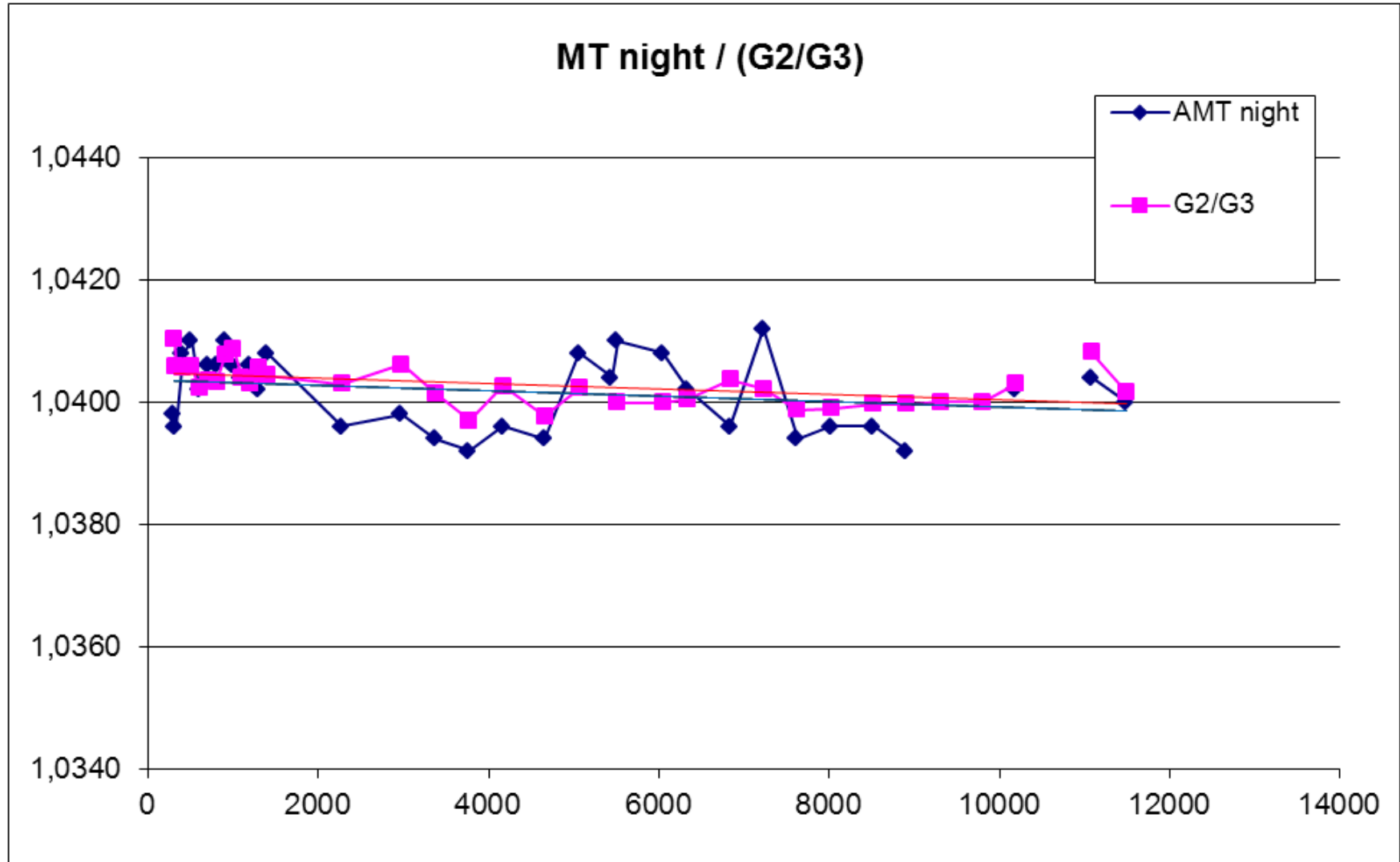
## MT Mode :

### MT cold day / MS day





## MT Mode :



**$A_{MT}(\text{night})$       constant**

**G2/G3                  constant**

**=> CALM stable, pas d'évolution relative gain onde longue**

**$A_{MS}$                   constant**

**$A_{MT}(\text{day})$         constant**

**=> pas d'évolution relative gain onde courte (sans filtre)**

**=> pas d'évolution relative gain onde courte (avec filtre)**

**=> pas d'évolution relative des filtres onde courte**

**=> Très bonne stabilité de l'instrument (filtre, canaux)**

The absolute calibration budget can be established :

L1A2		Bright clouds cold	Hot / bright scene	Night scene
		250 SW + 50 LW	210 SW + 130 LW (20°)	0 SW + 80 LW
Instrumental noise	Random	0.21%	0.11%	0.14%
Calibration CALM	Bias	0.12%	0.12%	0.12%
A' factor (0.2%)	Random	1%	0.3%	0%
<b>Registration and spectral effects</b>	Random	1.5%	0.55%	0%
Location	Random	0.4%	0.40%	0.40%
Budget @1σ %		<b>1.9%</b>	<b>0.8%</b>	<b>0.45%</b>
Budget @1σ W/m <sup>2</sup> /sr		<b>0.95</b>	<b>1.0</b>	<b>0.35</b>
Requirement			<b>1%</b>	

Items	Value	Type	
Short wave calibration (sphere)	3% @2 $\sigma$	Biais	1.5%
Error on spectral response		Biais	0.4%
Thermal gain correction	0.08%/°	Random	0.03%
Thermal leak correction	dT= 0.04° @1 $\sigma$ 20% of the thermal leak@1 $\sigma$	Random	0.04%
Location	0.06°@1 $\sigma$	Random	0.4%
Budget at 1 sigma			1.6%

**Absolute calibration of channel 2 was made in front of an integrating sphere at CNES facilities**

## SCARAB / Conclusion



**Radiometric noise**

**=> Very low, stable**

**Gain value**

**=> Very stable (0,05% per year)**

**A' factor**

**=> Very stable**